

# International Reference Ionosphere 2019

A COSPAR Capacity - Building Workshop  
Training Week, September 2 – 6, 2019

*Frederick University, Nicosia, Cyprus*



MINISTRY OF EDUCATION,  
YOUTH AND SPORTS



## **SUNDAY (1 September):**

Arrival

## **MONDAY (2 September):**

08.15-09:00 Registration

09:00-09:30 Opening Ceremony

09:30-10:30 Ionosphere-An introduction (B. Reinisch)

*10:30-11:00 Coffee/Tea*

11:00-12:00 IRI-Introduction and open problems (D. Bilitza)

12:00-12:30 IRIweb and related services (D. Bilitza)

*12:30-14:00 Lunch*

14:00-14:15 Forming teams and assigning research tasks

14:15-15:20 Team work with support from advisor and other lecturers

*15:20-15:40 Coffee/Tea*

15:40-17:00 Team work with support from advisor and other lecturers

## **TUESDAY (3 September):**

09:00-09:45 Ionosondes and the measurements they take (B. Reinisch)

09:45-10:30 GIRO and GAMBIT: access to ionosonde data (I. Galkin)

*10:30-11:00 Coffee/Tea*

11:00-11:45 ESA Space Weather Activities (A. Glover)

11:45-12:30 Space Weather applications: guidelines for users and developers (F. Da Dalt)

*12:30-14:00 Lunch*

14:15-15:20 Team work with support from advisor and other lecturers

*15:20-15:40 Coffee/Tea*

15:40-17:00 Team work with support from advisor and other lecturers

## **WEDNESDAY (4 September):**

09:00-10:00 GNSS data and ionospheric studies (A. Krankowski)

10:00-10:30 Access to GNSS data (A. Krankowski)

*10:30-11:00 Coffee/Tea*

11:00-11:45 Radio Occultation and access to COSMIC data (H. Haralambous)

11:45-12:30 IRTAM and Real-Time IRI (I. Galkin)

*12:30-14:00 Lunch*

14:15-15:20 Team work with support from advisor and other lecturers

*15:20-15:40 Coffee/Tea*

15:40-17:00 Team work with support from advisor and other lecturers

## **THURSDAY (5 September):**

09:00 -10:00 Coupling between ionosphere and thermosphere (S. Watanabe)

10:00 -10:30 Plasmasphere modelling (S. Watanabe)

*10:30-11:00 Coffee/Tea*

11:00-12:00 Incoherent scatter radar and ionospheric studies (S. R. Zhang)

12:00-12:30 Access to incoherent scatter data (S. R. Zhang)

*12:30-14:00 Lunch*

14:15-15:20 Team work with support from advisor and other lecturers

*15:20-15:40 Coffee/Tea*

15:40-17:00 Team work with support from advisor and other lecturers

## FRIDAY (6 September):

09:00-09:30 Science Talk: Real-time Modelling of the Vertical Total Electron Content (VTEC) Using B-splines (A. Goss et al.)

09:30-10:30 Representation of plasma temperatures in IRI (V. Truhlik)

*10:30-11:00 Coffee/Tea*

11:00 – 11:45 Representation of ion composition in IRI (V. Truhlik)

11:45 – 12:00 Dry run of student presentation – Team 1

12:00 – 12:15 Dry run of student presentation – Team 2

12:15 – 12:30 Dry run of student presentation – Team 3

*12:30-14:00 Lunch and Group-photo*

14:00 – 14:15 Dry run of student presentation – Team 4

14:15 – 14:30 Dry run of student presentation – Team 5

14:30 – 14:45 Dry run of student presentation – Team 6

14:45 – 15:00 Dry run of student presentation – Team 7

15:00 – 15:15 Dry run of student presentation – Team 8

*15:20-15:40 Coffee/Tea*

15:40-17:00 Team work with support from advisor and other lecturers

## List of Student Problems

**Problem A:** Investigate storm effects on foF2 at a location in the Northern hemisphere and one from a similar latitude region in the Southern hemisphere. What are the differences? Compare with IRI and IRI-Real-Time predictions. Use the 17-18 March 2015 (St. Patrick's Day) storm or select your own storm event. DATA: GIRO-Digisonde; MODEL: CCMC-IRIweb, Gambit-IRTAM; ADVISOR: I. Galkin

**Problem B:** Study the relationship between electron density ( $N_e$ ) and electron temperature ( $T_e$ ) in the middle ionosphere with Swarm and Incoherent Scatter data. Investigate how well IRI represents the relationship that you find between these parameters. IRI includes an option to use the anti-correlation between  $N_e$  and  $T_e$  to get near real-time  $T_e$  values if measured  $N_e$  values are available. Assess the reliability of this option based on your study results. DATA: Swarm, Madrigal-IS; MODEL: CCMC-IRIweb; ADVISOR: S. Watanabe

**Problem C:** Study the upper transition height. This is the height where the  $O^+$  density has dropped to 50% of the total ion density with the other 50% light ions ( $H^+$  and  $He^+$ ). Use Arecibo Incoherent Scatter (IS) data and compare with the prediction by IRI using the different topside ion density options. This parameter is not currently provided by IRI, so make suggestions how it could and should be included as an output parameter. DATA: Madrigal-IS; MODEL: CCMC-IRIweb; ADVISOR: S.-R. Zhang

**Problem D:** IRI-2016 includes three options for the height of the F2 peak, hmF2. Evaluate the different options with digisonde hmF2 measurements. Look into using also other data sources, like incoherent scatter radar observations or COSMIC radio occultation data. Determine which IRI option performs best for high and which for low solar activity. DATA: GIRO-Digisonde, COSMIC, Madrigal-IS; MODEL: CCMC-IRIweb; ADVISOR: D. Bilitza

**Problem E:** The Swarm satellites are measuring electron density and temperature. Help the ongoing assessment and validation of these data by comparing with IRI predictions using the different IRI options for these parameters at the satellite orbit altitudes. Explore existing validation studies and how much your results conform with these earlier results. Of particular interest is the Equatorial Ionization Anomaly region. DATA: Swarm; MODEL: CCMC-IRIweb; ADVISOR: V. Truhlik

**Problem F:** Study the ion temperature ( $T_i$ ) climatology in the middle ionosphere with incoherent scatter (IS) data. How does  $T_i$  change with altitude, latitude, and local time and how well is this variability represented by IRI. Investigate the availability of additional data sources for this parameter. DATA: Madrigal-IS; MODEL: CCMC-IRIweb; ADVISORS: S.-R. Zhang and V. Truhlik

**Problem G:** Study the extend of the Equatorial Ionization Anomaly (EIA) in Local Time with GPS TEC data. For what time period are separate cusps observed on both sides of the magnetic equator and when a single maximum at the magnetic equator. Compare with IRI predictions. DATA: IGS-TEC; MODEL: CCMC-IRIweb; ADVISOR: A. Krankowski

**Websites for getting data and model values:**

IRI model values using IRIweb at

[https://ccmc.gsfc.nasa.gov/modelweb/models/iri2016\\_vitmo.php](https://ccmc.gsfc.nasa.gov/modelweb/models/iri2016_vitmo.php).

(Questions: D. Bilitza)

Incoherent Scatter data from Madrigal (<http://madrigal.haystack.mit.edu/madrigal/> )

(Questions: S.-R. Zhang)

IRI real-time foF2 from the site <http://giro.uml.edu/IRTAM/> (Questions: I. Galkin)

Digisonde data from DIDBase at the Global Ionosphere Radio observatory (<http://giro.uml.edu>) (Questions: I. Galkin)

Swarm data from the Swarm data center at ESA at <https://earth.esa.int/web/guest/swarm/data-access> (Questions: V. Truhlik)

IGS-TEC data from NASA CDDIS archive at [https://cddis.nasa.gov/Data and Derived Products/](https://cddis.nasa.gov/Data_and_Derived_Products/) (Questions: A. Krankowski)

IGS-TEC movie display on NASA CDAWeb at <https://cdaweb.gsfc.nasa.gov/> (Questions: D. Bilitza)

COSMIC F-peak and topside electron densities for the COSMIC data center at <http://cdaac-www.cosmic.ucar.edu/cdaac/products.html> (Questions: H. Haralambous)

***A few additional notes:***

The first step is to develop a study plan with the different tasks to be accomplished and the assignments for different team members. Consult with the problem advisor on this step. Please contact the data/model source advisor (listed in the section “Websites for getting data and model values”) with questions you may have; he may have a faster way to get to the data that you need for your study. Please note that you are encouraged to use additional data sources if you have access to these data sources and if they are beneficial for solving your problem task. Plan on meeting as team every day and review and discuss the different tasks and how far team members have succeeded. Please keep plots and tables that you produce from early on for your final presentation. Towards the end of the first week sit down as team and discuss the layout of your final PowerPoint presentation. Plan on having about 10 slides maximum that include slides to cover the following important items: (1) state the problem, (2) explain your methodology to solve the problem, (3) explain the data and model used, and (4) your results and interpretation. Discuss with your problem advisor and benefit from the feedback that you receive during the dry runs on Friday of the first week.

## **Lecturers:**

**Dieter Bilitza** (George Mason University and NASA Goddard Space Flight Center, USA)

**Bodo Reinisch** (University of Massachusetts Lowell and Digisonde International Inc., USA)

**Ivan Galkin** (University of Massachusetts Lowell, USA)

**Vladimir Truhlik** (Institute of Atmospheric Physics, Prague, Czech Republic)

**Shigeto Watanabe** (Hokkaido University, Sapporo, Japan)

**Shunrong Zhang** (MIT, Millstone Hill, USA)

**Andrzej Krankowski** (University of Warmia and Mazury, Olsztyn, Poland)

**Alexi Glover** (ESA, ESOC, Darmstadt, Germany)

**Federico Da Dalt** (Rhea for ESA, Darmstadt, Germany)

**Haris Haralambous** (Fredrick University, Nicosia, Cyprus)